LoCo Lab EDGES Memo 189 Alternatives for Beam Correction

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1 Introduction

The antenna temperature spectrum for a sky temperature T_{sky} measured with an achromatic antenna beam $B(\Omega, \nu_0)$ is given by

$$T(\nu) = \int T_{sky}(\Omega, \nu) B(\Omega, \nu_0) d\Omega.$$
(1)

In general, however, the beam is chromatic and the measured spectrum is given by

$$T_m(\nu) = \int T_{sky}(\Omega, \nu) B(\Omega, \nu) d\Omega.$$
(2)

A spectrum corrected for beam chromaticity can be obtained from the measured spectrum by dividing the measured spectrum by the chromaticity correction factor C,

$$C(\nu) = \frac{T_m^*(\nu)}{T^*(\nu)},$$
(3)

where

$$T^*(\nu) = \int T^*_{sky}(\Omega,\nu)B^*(\Omega,\nu_0)d\Omega$$
(4)

and

$$T_m^*(\nu) = \int T_{sky}^*(\Omega, \nu) B^*(\Omega, \nu) d\Omega$$
(5)

are simulated spectra for an achromatic and chromatic beam, respectively. They are computed using the models T^*_{sky} and B^* for the sky temperature and antenna beam, respectively.

2 Beam Correction Alternatives

Two possible beam corrections approaches are presented for consideration when integrating measurements across time (LST or GHA). Here, t1, t2, t3, ... represent the time stamps of the raw spectra and N_t is the total number of raw spectra.

2.1 Alternative 1: Time-averaging the corrected spectra

$$\left\langle \frac{T_m}{C} \right\rangle_t = \frac{1}{N_t} \left[\frac{T_m(t1)}{C(t1)} + \frac{T_m(t2)}{C(t2)} + \frac{T_m(t3)}{C(t3)} \dots \right]$$
(6)

2.2 Alternative 2: Correcting the time-average of uncorrected spectra

$$\frac{\langle T_m \rangle_t}{\langle T_m^* \rangle_t} = \frac{\frac{1}{N_t} \left[T_m(t1) + T_m(t2) + T_m(t3) \dots \right]}{\frac{\frac{1}{N_t} \left[T_m^*(t1) + T_m^*(t2) + T_m^*(t3) \dots \right]}{\frac{1}{N_t} \left[T^*(t1) + T^*(t2) + T^*(t3) \dots \right]}}$$
(7)

Both alternatives approach the following average

$$\frac{1}{N_t} \left[T(t1) + T(t2) + T(t3) \dots \right]$$
(8)

as the models for the sky temperature and beam approach the true values.